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# Tracing Earth's O<sub>2</sub> evolution using Zn/Fe ratios in marine carbonates

Xiao-Ming Liu<sup>1,2\*</sup>, Linda C. Kah<sup>3</sup>, Andrew H. Knoll<sup>4</sup>, Huan Cui<sup>5</sup>,

Alan J. Kaufman<sup>5</sup>, Anat Shahr<sup>2</sup>, and Robert M. Hazen<sup>2</sup>

1. *Department of Geological Sciences, University of North Carolina, Chapel Hill, North Carolina, 27599 USA*
2. *Geophysical Laboratory, Carnegie Institution of Washington, Washington DC 20015, USA*
3. *Department of Earth and Planetary Sciences, University of Tennessee, Knoxville, Tennessee 37996, USA*
4. *Department of Organismic and Evolutionary Biology, Harvard University, Cambridge, Massachusetts 02138, USA*
5. *Department of Geology, University of Maryland, College Park, Maryland 20742, USA*

\* Corresponding author:

Department of Geological Sciences  
University of North Carolina  
Chapel Hill, North Carolina, 27599 USA

Tel: (919) 926-0675  
Email: [xiaomliu@unc.edu](mailto:xiaomliu@unc.edu)

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## 1   **Abstract**

2   Through Earth history, atmospheric oxygen has increased from initial values near zero to its  
3   present day level of about 21% by volume; concomitantly, changes in ocean redox conditions  
4   have fundamentally altered global biogeochemical cycles. While there is a reasonable  
5   understanding of where oxygen history begins and ends, the quantitative timetable of  
6   oxygenation that links the endpoints has proven contentious. Equilibrium between marine  
7   surface environments and the overlying atmosphere suggests that carbonate-based redox  
8   proxies could refine paleoredox records in time and space. Here we explore the use of Zn/Fe  
9   ratios to infer the evolution of atmospheric O<sub>2</sub> through time, based on marine carbonate  
10   rocks that are well characterized in terms of depositional age, environmental setting, and  
11   diagenetic history. While Fe and Zn in the shallow ocean is mainly sourced from  
12   hydrothermal inputs, their redox sensitivities differ significantly, so that geological intervals  
13   with higher O<sub>2</sub> would be characterized by stepped increases in Zn/Fe as preserved in shallow  
14   marine carbonates. Therefore, Zn/Fe analyses of ancient carbonates allow us to constrain past  
15   atmospheric pO<sub>2</sub> levels, providing a secular record of atmospheric O<sub>2</sub> over the past 3.5  
16   billion years. In particular, we corroborate an earlier proposal that for much of the  
17   Proterozoic Eon, O<sub>2</sub> levels were as low as 0.1-1% of present atmospheric level. We conclude  
18   that Zn/Fe in shallow marine carbonate rocks has potential to provide a quantitative tracer for  
19   the long-term redox evolution of the oceans and the rise of atmospheric O<sub>2</sub>.

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## Introduction

Earth's O<sub>2</sub>-rich atmosphere, unique among known planets, has played an essential role in evolving feedbacks between life and environment. Atmospheric O<sub>2</sub> was extremely low in the Archean Eon (>2.5 billion years ago), and while multiple lines of evidence suggest that Earth's oxygenation was protracted (Kah and Bartley, 2011; Kah et al., 2004; Lyons et al., 2014; Planavsky et al., 2014), pO<sub>2</sub> may have risen abruptly at two different points in time: first during the "Great Oxygenation Event" (GOE) at ~2.4 billion years (Ga) (Canfield, 2005; Farquhar et al., 2011; Guo et al., 2009; Holland, 2006), when atmospheric O<sub>2</sub> rose from <0.001% to an intermediate value commonly estimated as 1 to 10% of the current level (Farquhar et al., 2000; Pavlov and Kasting, 2002), and again during a "Neoproterozoic Oxygenation Event" (NOE) at ~800 to 542 million years ago (Canfield and Teske, 1996; Fike et al., 2006; Frei et al., 2009; Och and Shields-Zhou, 2012). The latter transition may well have continued into the Phanerozoic Eon, eventually resulting in near-present O<sub>2</sub> (Berner, 2006; Dahl et al., 2010; Sperling et al., 2015).

Redox-sensitive major and trace elements in iron formations and black shales deposited beneath euxinic waters have been developed as proxies to reconstruct paleoenvironmental history in deep time (Konhauser et al., 2009; Sahoo et al., 2012; Scott et al., 2008). The paucity of these facies in many Proterozoic successions, however, limits the continuity of current reconstructions of Earth's oxygenation. Here we provide evidence for the hypothesis that carbonate-based redox proxies can provide an independent estimate of past pO<sub>2</sub>, expanding the paleo-redox record in time and space (Hardisty et al., 2014). Limestone and penecontemporaneous dolomites that retain depositional signatures well (Wilson et al., 2010) are abundant in the geologic record, typically recording shallow marine

environments that would have been in open communication with the overlying atmosphere. Paleoenvironmental research on carbonate rocks commonly focuses on individual stratigraphic successions; here we adopt a complementary strategy, analyzing a large suite of Phanerozoic, Proterozoic, and Archean samples that enables us to make statistical statements (Sperling et al., 2015) about Zn/Fe in the global surface ocean through geologic time. More importantly, we develop a new tool to provide quantitative constraints on atmospheric  $pO_2$  through Earth history.

In the modern ocean, zinc input from hydrothermal ridge systems ( $\sim 4.4 \times 10^9$  mol/year) is an order of magnitude greater than riverine fluxes ( $\sim 3.4 \times 10^8$  mol/year) (Robbins et al., 2013). As an essential nutrient in many phytoplankton enzymes, especially those of eukaryotes (Williams and da Silva, 1996), zinc plays an important role in marine primary production, and for this reason, Zn is depleted in surface waters relative to the deep sea (Morel and Price, 2003). Zn concentrations in euxinic black shale and iron formations (Robbins et al., 2013; Scott et al., 2013), however, suggest that the bioavailability of Zn has not changed dramatically through Earth history. The Fe budget is similar to that of Zn, wherein hydrothermal input dominates over riverine fluxes by a factor of  $\sim 9$  (Wheat et al., 2002)(Wheat et al., 2002)(Wheat et al., 2002)(Wheat et al., 2002)(Wheat et al., 2002)(Wheat et al., 2002). Under sulphidic conditions, dissolved  $Zn^{2+}$  and  $Fe^{2+}$  behave similarly and are rapidly precipitated as sulphides (Morse and Luther III, 1999). In addition, because both Fe and Zn behave as incompatible elements during mantle partial melting, Zn/Fe has been developed as a tracer of mantle redox, revealing that the oxygen fugacity of the upper mantle has remained relatively constant through Earth history (Lee et al., 2010). In the following discussion, we assume that Zn/Fe in hydrothermal inputs into the ocean have not changed

significantly through time. We recognize, however, that a number of factors could limit this assumption, and consider these below.

Zn/Fe in the sedimentary record thus has the potential to document Earth surface redox evolution if we consider the following assumptions: 1) Zn and Fe budgets in oceans are dominated by hydrothermal inputs and are therefore not significantly influenced by secular evolution of continental inputs; 2)  $\text{Fe}^{2+}$  and  $\text{Zn}^{2+}$  have similar solubility in oceans; 3) the partition coefficient of Zn/Fe ratios into carbonates have remained the same through time; and 4) when  $\text{Fe}^{2+}$  is oxidized to  $\text{Fe}^{3+}$ , it precipitates from seawater and thus is not incorporated into carbonate; zinc, however, remains divalent as  $\text{Zn}^{2+}$ .

## Methods

Major, trace, and REE concentrations were determined with a Thermo Scientific® iCAP-Q ICP-MS (Inductively Coupled Plasma – Mass Spectrometry) at the Carnegie Institution of Washington. Approximately 5 to 10 mg of micro-drilled sample powders were weighed and dissolved in 2 ml distilled 0.4 M  $\text{HNO}_3$  and reacted for 12 hours. The resulting solutions were centrifuged for 5 minutes at ~6000 rps and 1 ml of the supernatant was pipetted and diluted with distilled 4 ml 0.4 M  $\text{HNO}_3$  for elemental analysis. Calibration curves were created using multi-elemental standards with different dilution made from pure element solutions (Alfa Aesar®). Both standard and sample solutions were doped with 4 ppb In to correct for instrumental drift. Precision of the analyses was determined by repeated analyses of an in-house carbonate standard, and was typically better than 5% ( $2\sigma$ ) for major elements, and better than 10% ( $2\sigma$ ) for most trace elements including rare earth

elements (REEs). Accuracy of the analyses was determined by replicates of an international coral standard (JCp-1), as shown in Fig. S1.

## **Results and discussion**

Here, we report Zn/Fe molar ratios in marine carbonate rock through Earth history (Fig. 1) and provide a quantification of atmospheric O<sub>2</sub> evolution since the Mesoarchean Era. Samples (n = 1700) come from our analyses (n = 300), as well as a literature compilation (see Table S1). In all carbonate samples, the potential for diagenetic alteration is of concern. To evaluate the degree of sample alteration, we selected specimens with known sedimentological and stratigraphic context and investigated their petrography and elemental and isotope geochemistry. Samples used in this study were primarily composed of fine-grained limestone and penecontemporaneous, fabric-retentive dolostone, including both micrites and stromatolites. We micro-sampled carbonate specimens from polished billets to avoid weathering alteration, secondary veins/precipitation, and areas with visible non-carbonate phases. In addition to geological and petrographic criteria, we further selected samples based on primary isotopic and trace element patterns (see SI 2).

Even if we carefully select the most primary samples, we cannot ignore diagenetic influences on the elemental composition of sampled carbonates, as this can contribute to local variation of Zn/Fe (Fig. 1). Both Zn and Fe partition coefficient ( $K_d$ ) from fluid to carbonates increase with increasing diagenesis as shown by earlier work of Brand and Veizer (1980). Also,  $K_d$  (Fe) increases faster compared to  $K_d$  (Zn), from 1 to 20 and from 5.2 to 5.5 for Fe and Zn, respectively. According to this work, diagenesis will cause decrease in Zn/Fe ratios by incorporating more Fe than Zn in carbonates. We acknowledge that all of the

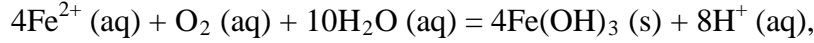
carbonates examined here have undergone some degree of burial diagenesis, and this will be reflected in the variance of Zn/Fe within individual time intervals. Also, local primary production differences may contribute to Zn/Fe variability of different formations from the same interval. In the modern oxidized shallow ocean, particulate Fe sourced from eroding continents remains biogeochemically labile and may be cycled back to a dissolved phase during diagenesis in reducing continental margin sediments (Raiswell et al., 2006). Therefore, there is also a potentially large and variable source of reactive Fe to shallow marine settings that is decoupled from the Zn flux, which likely cause Zn/Fe ratios to be lower and therefore could contribute to the variations observed in Zn/Fe data. In addition, theoretical calculations suggest that kinetic effects on trace element partitioning in carbonate may contribute to Zn/Fe variability in samples from the same locality (DePaolo, 2011; Watson, 2004). Importantly, however, these influences should not result in systematic variations that would contribute to observed first-order secular changes. We plot all carbonate samples based on their lithology in Fig. S2; this shows that there is no systematic difference between limestone and dolomite samples through time -- not unexpected, as many Proterozoic dolomites formed penecontemporaneously and preserve geochemical signatures as well as coeval limestones that underwent neomorphism during burial.

We observe a distinct trend of increasing Zn/Fe through time (Fig. 1), especially around the GOE and NOE. Our Paleoproterozoic data are also consistent with earlier suggestions that  $pO_2$  may have risen substantially during the GOE and then declined again to persistent Proterozoic values (Lyons et al., 2014). Employing three statistically complementary approaches (see details in SI 3), carbonate Zn/Fe could follow “step” or “smooth” fits through Earth’s history (Figs. S3, 4, & 5), where we prefer the “step” approach



with lognormal distributions (see Fig. 2 and SI 3). Using lognormal distributions to estimate Zn/Fe through time, we can provide quantitative constraints on Earth's atmospheric O<sub>2</sub> evolution, as follows.

From the chemical reaction of Fe oxidization from Fe<sup>2+</sup> to Fe<sup>3+</sup>:



$$K = \frac{a_{(\text{H}^+)}^8}{a_{(\text{H}_2\text{O})}^{10} \times a_{(\text{Fe}^{2+})}^4 \times a_{(\text{O}_2)}} \quad [1],$$

where  $K$  is the equilibrium constant and  $a$  is activity. In this equation, we assume that when Fe<sup>2+</sup> oxidizes to Fe<sup>3+</sup> and is precipitated from the aqueous system as iron hydroxide, only Fe<sup>2+</sup> gets incorporated into carbonates. We are aware that secular variations in seawater sulfate might modulate hydrothermal iron fluxes through time via the formation of iron sulfides (Kump and Seyfried, 2005), we do not know the extent to which Zn abundances might similarly be buffered and so do consider this in our first-order model. Assuming O<sub>2</sub> equilibrium between atmosphere and surface ocean on hundred million year time scales, we can write the equation using atmospheric oxygen fugacity,  $f_{(\text{O}_2)}$ , as

$$\log K = 4\log \frac{1}{a_{(\text{Fe}^{2+})}} - \log f_{(\text{O}_2)} + \log \frac{a_{(\text{H}^+)}^8}{a_{(\text{H}_2\text{O})}^{10}} \quad [2],$$

if we assume the Zn concentrations in seawater and partitioning of Zn/Fe from seawater to carbonate minerals are constant over Earth history. Therefore, we can write the equation normalized to Zn<sup>2+</sup> as

$$\log K = 4\log a_{(\text{Zn}^{2+}/\text{Fe}^{2+})}^P - \log f_{(\text{O}_2)}^P + \log \frac{a_{(\text{H}^+)}^8}{a_{(\text{H}_2\text{O})}^{10}}$$

$$\log K = 4 \log a_{(Zn^{2+}/Fe^{2+})}^M - \log f_{(O_2)}^M + \log \frac{a_{(H^+)}^8}{a_{(H_2O)}^{10}} \quad [3],$$

in which superscripts P and M indicate the past and modern parameters. Assuming pH and K are constant (see SI 4), we can simplify the relationship between Fe/Zn ratios and  $f_{(O_2)}$ , as  $f_{(O_2)}^P = r^4 \cdot f_{(O_2)}^M$  [4], where  $f_{(O_2)}^P$  is the oxygen fugacity in the past (any time in Earth's history),  $f_{(O_2)}^M$  is the oxygen fugacity in modern time, and

$$r = \frac{(Zn^{2+}/Fe^{2+})^P}{(Zn^{2+}/Fe^{2+})^M} \quad [5],$$

provides Zn/Fe ratios in past carbonate normalized to modern values. If we assume that atmospheric  $O_2$  is in equilibrium with the shallow marine environment, and that we know the current atmospheric  $P_{(O_2)}$  (0.21) and the modern seawater Zn/Fe ratios as reflected in Zn/Fe ratios of marine carbonates, we can use Zn/Fe to calculate  $f_{(O_2)}$  at any given time of Earth history (Fig. 3). This  $f_{(O_2)}$  curve provides a more continuous coverage of atmospheric  $O_2$  levels compared to compilations derived from multiple geochemical tracers, such as mass-independent S isotopes and paleosol records (Catling and Claire, 2005; Rye and Holland, 1998).

The  $\log f_{(O_2)}$  curve in Fig. 3 reproduces what we think we know about oxygen history: estimated  $pO_2$  is extremely low in the Archean and reaches modern levels only in the mid-Paleozoic Era. Moreover, the estimates match our current understanding (Lyons et al., 2014) of a general two-step increase of atmospheric  $O_2$  around the GOE and the NOE. Importantly, our study provides an estimate of the upper and lower bounds on  $pO_2$  in the mid-Proterozoic atmosphere, with a preferred value between 0.1 and 1% PAL. This value is substantially lower than traditional estimates based on paleosol work (Canfield, 1998; Rye and Holland,

1998), but consistent with recent estimates based on an independent tracer, a kinetic model for Cr-Mn oxidation and Cr isotopes in ironstones (Planavsky et al., 2014). We conducted sensitivity tests of temperature and pH variations on our  $pO_2$  estimates and found that the influence of temperature is negligible. pH, however, could potentially lower  $pO_2$  estimates, especially for earlier samples when  $pCO_2$  was high (see SI 4 for details); thus, our estimates of Proterozoic  $pO_2$  should be considered conservative and may overestimate past oxygen levels. There are hints of biologically interesting structure in the Neoproterozoic and Cambrian records, but at present our sample numbers and bin sizes are too small to address this in detail. As more carbonate data become available for key transitional time periods such as those around GOE and NOE, potentially complex secular patterns of redox change may become clearer. Further investigations on well-constrained modern and Phanerozoic marine carbonates are currently underway to evaluate with more quantitative rigor the potential effects of diagenesis, mineralogy, and ocean depth gradients distributions on the proxy proposed here.

## **Conclusion**

In summary, we have demonstrated the potential for using divalent cations in carbonates as sensitive proxies for the evolution of Earth's near surface environment. Because many marine carbonate rocks were deposited in shallow marine environments, in direct contact with the atmosphere, elemental ratios are likely to reflect equilibrium atmospheric conditions extending back to the Archean Eon and including time intervals poorly represented by other lithologies. Although further work will be needed to fully validate this promising paleo-redox proxy, carbonate-based redox proxies show great

- 197 potential to expand the paleoredox record and to provide self-consistent and quantitative
- 198 constraints on atmospheric O<sub>2</sub> through Earth's history.

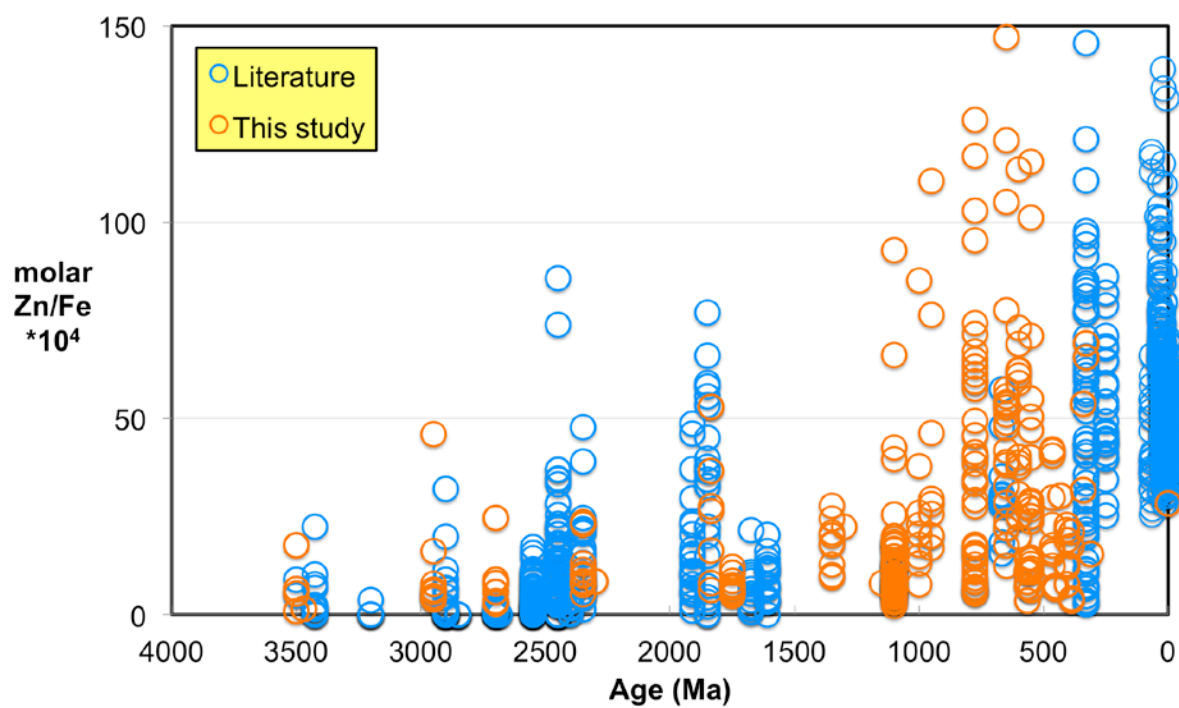
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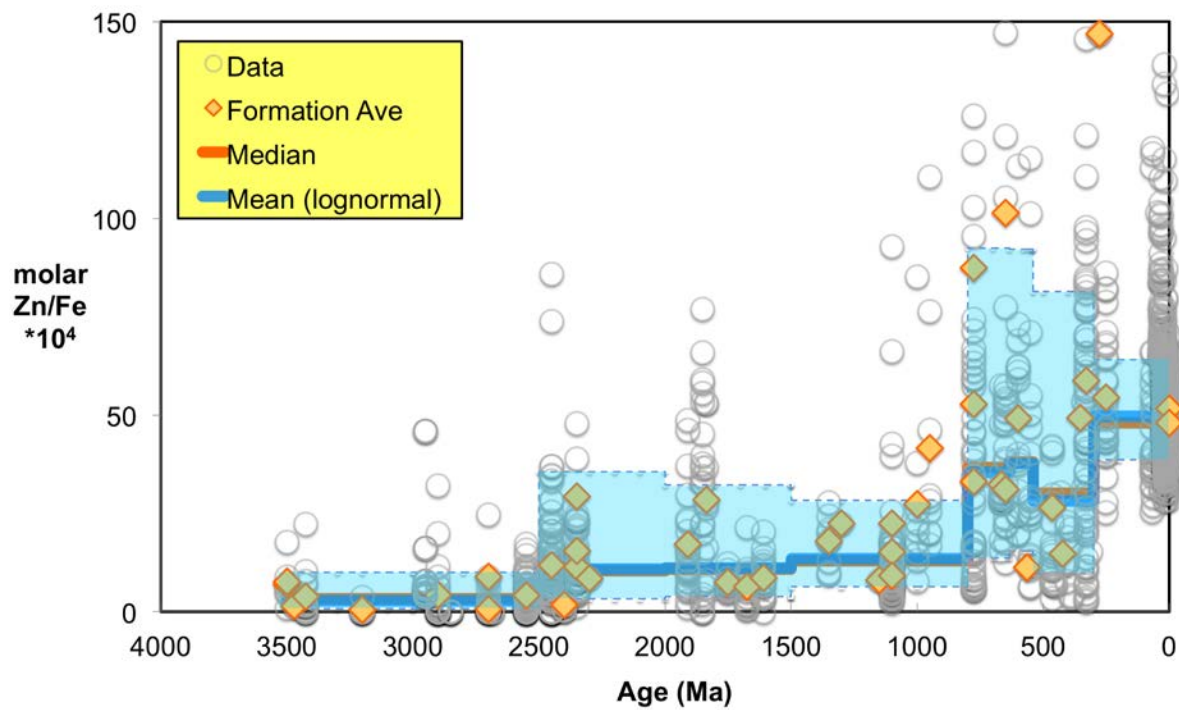
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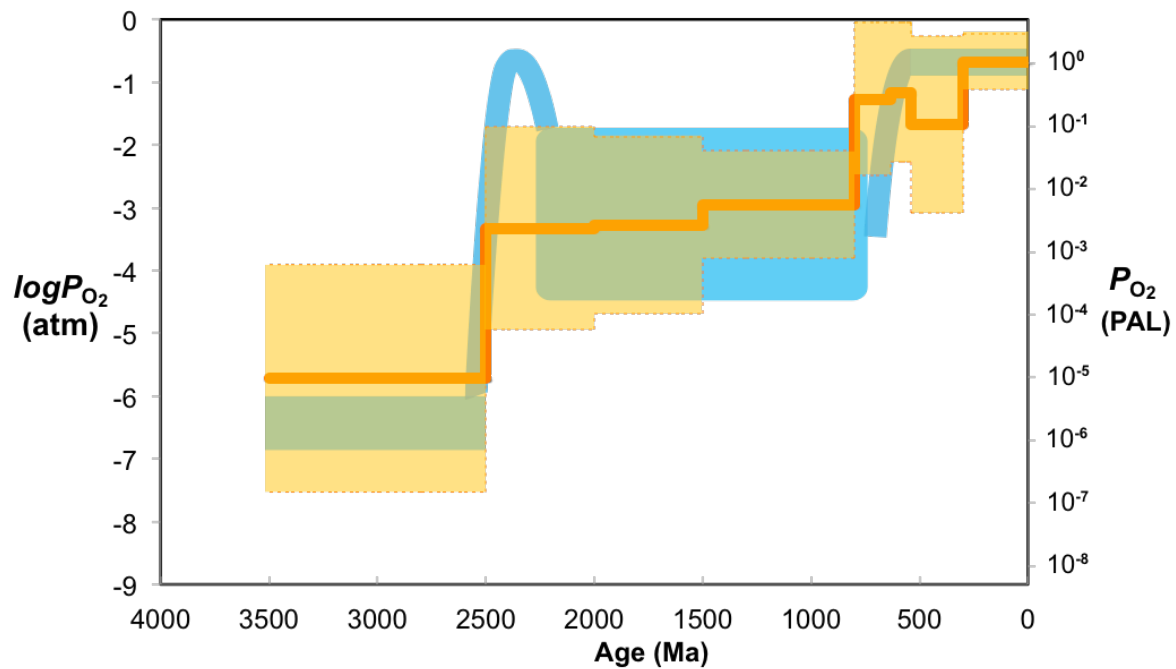
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220 Figure 1.



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222 Figure 2.



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224 Figure 3.



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322

## 323 **Author Contributions**

324   X-ML and RMH designed the project with inputs from all authors. X-ML performed the  
325   chemical analyses. X-ML wrote the manuscript with inputs from all authors. LK, AHK, HC,  
326   AJK, and RMH provided samples.

327

## 328 **Competing financial interests**

329   The authors declare no competing financial interests.

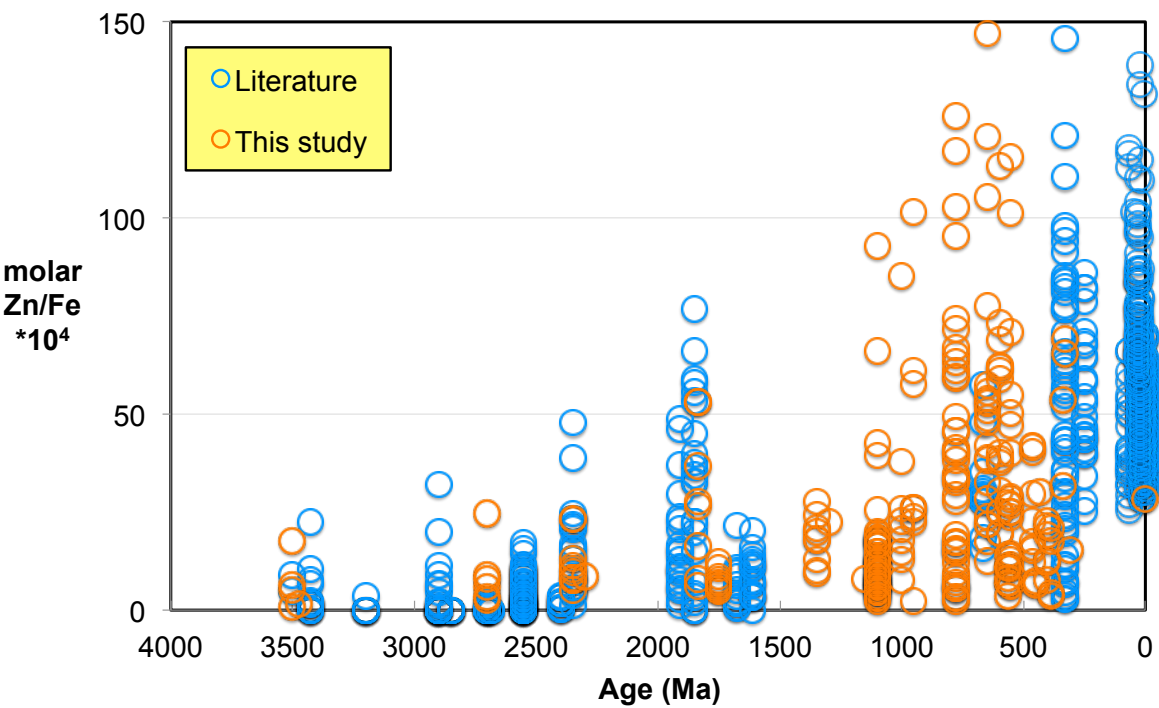
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155

156 Figure 1.

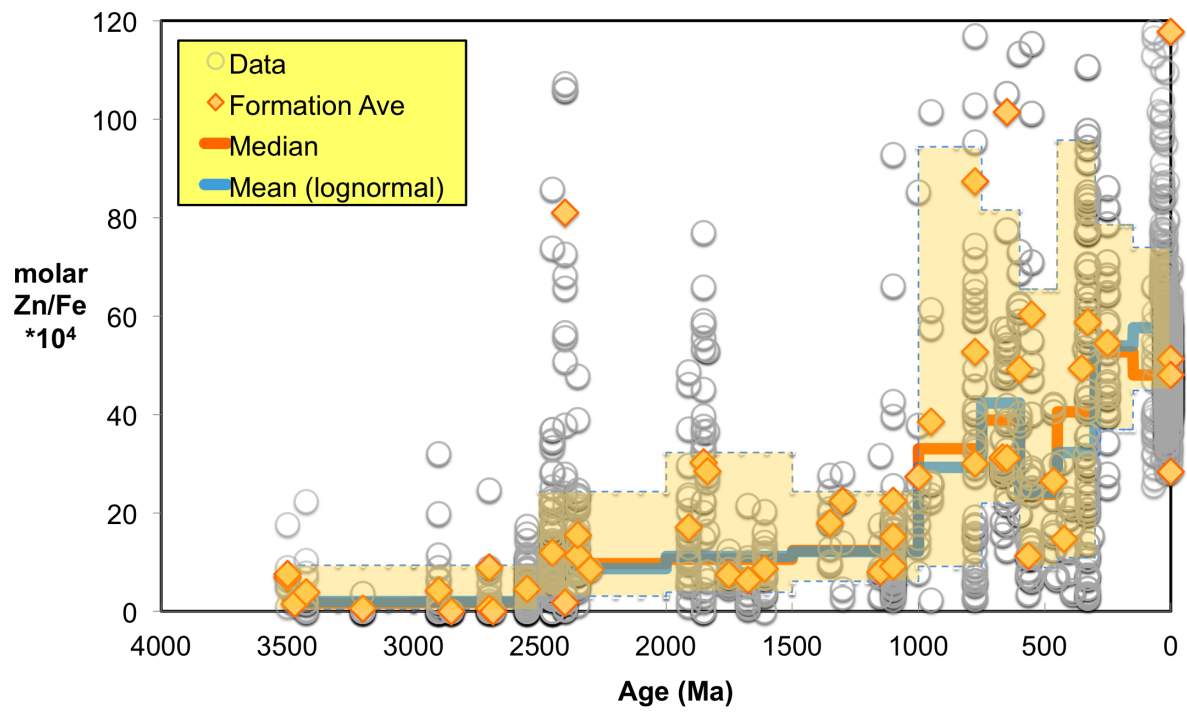
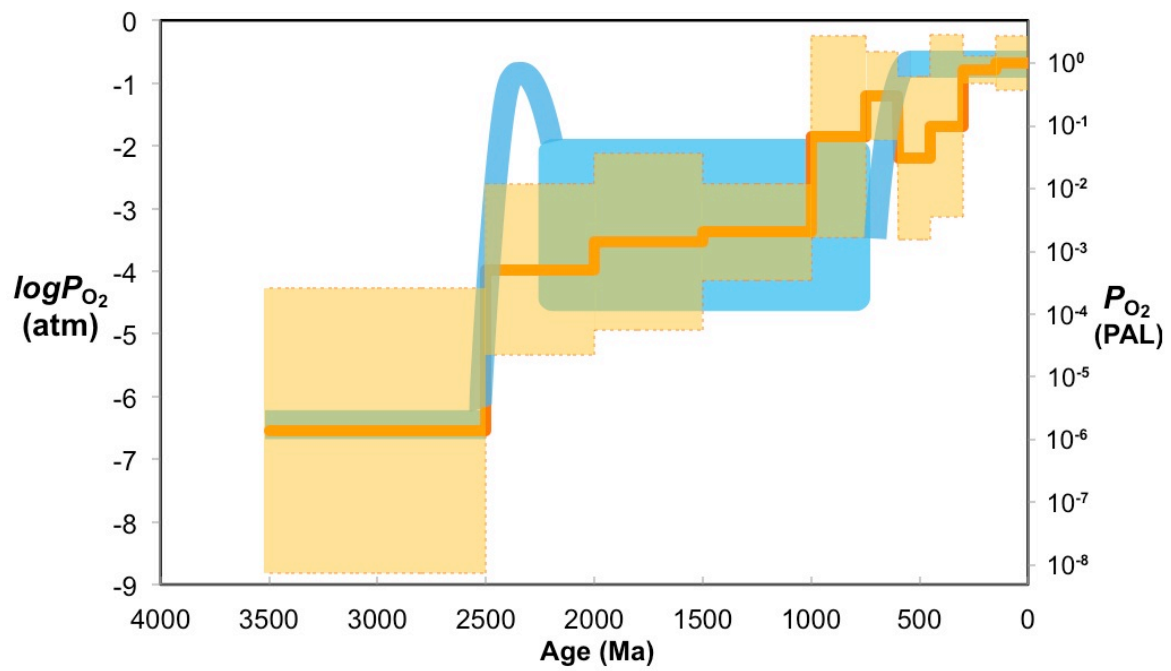


Figure 2.



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161 Figure 3.

## Supplementary Information

### SI 1. Tables

**Table S1. Zn/Fe ratios with sample name and age information from this study.**

Geologic Unit	Sample #	Age (Ma)	Zn/Fe *10 <sup>4</sup>	Reference
Strelley Pool Fm, Warrawoona Group, Pilbara, Western Australia	BH2	3400	0.9	Hazen collection
Strelley Pool Fm, Warrawoona Group, Pilbara, Western Australia	BH3	3400	17.6	Hazen collection
Strelley Pool Fm, Warrawoona Group, Pilbara, Western Australia	BH7	3400	5.3	Hazen collection
Calcite filling in Mt. Ada basalt	BH18	3470	1.5	Hazen collection
Pongola Fm, South Africa	Po1	2950	16.1	Beukes and Lowe 1989
Pongola Fm, South Africa	Po2	2950	45.9	Beukes and Lowe 1989
Pongola Fm, South Africa	Po3	2950	4.5	Beukes and Lowe 1989
Pongola Fm, South Africa	Po4	2950	6.4	Beukes and Lowe 1989
Pongola Fm, South Africa	Po5	2950	5.9	Beukes and Lowe 1989
Pongola Fm, South Africa	Po6	2950	7.8	Beukes and Lowe 1989
Pongola Fm, South Africa	Po7	2950	5.8	Beukes and Lowe 1989
Pongola Fm, South Africa	Po8	2950	4.0	Beukes and Lowe 1989
Pongola Fm, South Africa	Po9	2950	3.8	Beukes and Lowe 1989
Pongola Fm, South Africa	Po10	2950	4.8	Beukes and Lowe 1989
Tumbiana Fm, Fortesue Group, Western Australia	BH10	2700	6.5	Hazen collection
Tumbiana Fm, Fortesue Group, Western Australia	BH12	2700	24.6	Hazen collection
Tumbiana Fm, Fortesue Group, Western Australia	BH13	2700	8.9	Hazen collection
Tumbiana Fm, Fortesue Group, Western Australia	BH14	2700	8.0	Hazen collection
Tumbiana Fm, Fortesue Group, Western Australia	BH15	2700	2.8	Hazen collection
Tumbiana Fm, Fortesue Group, Western Australia	BH16	2700	3.4	Hazen collection
Kazput Fm, Turee Creek, Australia	AN94	2350	7.6	Knoll et al. unpublished
Kazput Fm, Turee Creek, Australia	AN95	2350	8.4	Knoll et al. unpublished
Kazput Fm, Turee Creek, Australia	AN96	2350	9.9	Knoll et al. unpublished
Kazput Fm, Turee Creek, Australia	AN97	2350	7.7	Knoll et al. unpublished
Kazput Fm, Turee Creek, Australia	AN98	2350	8.2	Knoll et al. unpublished
Kazput Fm, Turee Creek, Australia	AN99	2350	5.2	Knoll et al. unpublished
Kazput Fm, Turee Creek, Australia	AN100	2350	7.8	Knoll et al. unpublished
Kazput Fm, Turee Creek, Australia	AN101	2350	10.4	Knoll et al. unpublished
Kazput Fm, Turee Creek, Australia	AN102	2350	11.5	Knoll et al. unpublished
Kazput Fm, Turee Creek, Australia	AN103	2350	13.1	Knoll et al. unpublished
Kazput Fm, Turee Creek, Australia	AN104	2350	23.0	Knoll et al. unpublished
Kazput Fm, Turee Creek, Australia	AN105	2350	23.5	Knoll et al. unpublished
Kazput Fm, Turee Creek, Australia	BH19	2350	8.5	Hazen collection
Duck Creek, Australia	AN80	1835	53.0	Wilson et al. 2010
Duck Creek, Australia	AN82	1835	8.0	Wilson et al. 2010
Duck Creek, Australia	AN83	1835	36.6	Wilson et al. 2010



Duck Creek, Australia	AN84	1835	53.0	Wilson et al. 2010
Duck Creek, Australia	AN85	1835	26.5	Wilson et al. 2010
Duck Creek, Australia	AN86	1835	16.2	Wilson et al. 2010
Duck Creek, Australia	AN87	1835	27.6	Wilson et al. 2010
Duck Creek, Australia	AN88	1835	6.2	Wilson et al. 2010
Vempalle Fm, Cuddapah Basin, India	V-SC/1	1750	8.5	Chakrabarti et al. 2013
Vempalle Fm, Cuddapah Basin, India	V-SC/2	1750	4.8	Chakrabarti et al. 2013
Vempalle Fm, Cuddapah Basin, India	V-SC/3	1750	5.3	Chakrabarti et al. 2013
Vempalle Fm, Cuddapah Basin, India	V-SC/5	1750	10.6	Chakrabarti et al. 2013
Vempalle Fm, Cuddapah Basin, India	V-SC/6-1	1750	6.8	Chakrabarti et al. 2013
Vempalle Fm, Cuddapah Basin, India	V-P2/1	1750	5.5	Chakrabarti et al. 2013
Vempalle Fm, Cuddapah Basin, India	V-P8/1	1750	12.2	Chakrabarti et al. 2013
Vempalle Fm, Cuddapah Basin, India	V-P10/1	1750	7.5	Chakrabarti et al. 2013
Vempalle Fm, Cuddapah Basin, India	V-P11/1	1750	6.9	Chakrabarti et al. 2013
Vempalle Fm, Cuddapah Basin, India	V-P12/1	1750	6.1	Chakrabarti et al. 2013
Kyrpy Group, East European Platform, Southern Urals, Russia	C133-3252.3	1350	17.8	Kah et al. 2007
Kyrpy Group, East European Platform, Southern Urals, Russia	C133-3253.3	1350	12.8	Kah et al. 2007
Kyrpy Group, East European Platform, Southern Urals, Russia	C133-3038	1350	9.5	Kah et al. 2007
Kyrpy Group, East European Platform, Southern Urals, Russia	C133-2767.5	1350	27.7	Kah et al. 2007
Kyrpy Group, East European Platform, Southern Urals, Russia	C133-3114.5	1350	20.1	Kah et al. 2007
Kyrpy Group, East European Platform, Southern Urals, Russia	C203-3852A	1350	24.4	Kah et al. 2007
Kyrpy Group, East European Platform, Southern Urals, Russia	C203-2459.8	1350	10.0	Kah et al. 2007
Kyrpy Group, East European Platform, Southern Urals, Russia	C203-2459.6-1	1350	18.1	Kah et al. 2007
Kyrpy Group, East European Platform, Southern Urals, Russia	C203-2353	1350	20.9	Kah et al. 2007
Sulky Fm, Dismal Lakes, Canada	DL1-364-1	1300	13.9	Kah et al. 2006
Sulky Fm, Dismal Lakes, Canada	DL1-306-1	1300	28.0	Kah et al. 2006
Sulky Fm, Dismal Lakes, Canada	DL1-332	1300	22.4	Kah et al. 2006
Sulky Fm, Dismal Lakes, Canada	SL16-1-1	1300	2.7	Kah et al. 2006
Sulky Fm, Dismal Lakes, Canada	SL17-10-1	1300	4.5	Kah et al. 2006
Avzyan Fm, Southern Urals, Russia	M1(AZ)-39	1150	3.0	Bartley et al. 2007
Avzyan Fm, Southern Urals, Russia	M1(AZ)-47	1150	31.7	Bartley et al. 2007
Avzyan Fm, Southern Urals, Russia	RV(AZ)-15	1150	8.8	Bartley et al. 2007
Avzyan Fm, Southern Urals, Russia	RV(AZ)-33	1150	13.8	Bartley et al. 2007
Avzyan Fm, Southern Urals, Russia	KT(AZ)-49.5	1150	17.6	Bartley et al. 2007
Avzyan Fm, Southern Urals, Russia	KT(AZ)-131.6	1150	3.0	Bartley et al. 2007
Avzyan Fm, Southern Urals, Russia	KT(AZ)-236-1	1150	7.0	Bartley et al. 2007
Avzyan Fm, Southern Urals, Russia	KT(AZ)-373.5	1150	8.0	Bartley et al. 2007
El Mreiti, Atar Group, West Africa	F4-10-1	1100	42.6	Gilleaudeau and Kah 2013
El Mreiti, Atar Group, West Africa	F4-19-1	1100	17.7	Gilleaudeau and Kah 2013
El Mreiti, Atar Group, West Africa	F4-50-1	1100	13.9	Gilleaudeau and Kah 2013
El Mreiti, Atar Group, West Africa	F4-53-1	1100	16.4	Gilleaudeau and Kah 2013
El Mreiti, Atar Group, West Africa	F4-90-1	1100	17.8	Gilleaudeau and Kah 2013
El Mreiti, Atar Group, West Africa	F4-95-1	1100	11.9	Gilleaudeau and Kah 2013
El Mreiti, Atar Group, West Africa	F4-98-1	1100	17.7	Gilleaudeau and Kah 2013
El Mreiti, Atar Group, West Africa	F4-99-1	1100	39.6	Gilleaudeau and Kah 2013

El Mreiti,Atar Group, West Africa	F4-102-1	1100	20.0	Gilleaudeau and Kah 2013
El Mreiti,Atar Group, West Africa	F4-104-1	1100	8.6	Gilleaudeau and Kah 2013
El Mreiti,Atar Group, West Africa	F4-106-1	1100	11.4	Gilleaudeau and Kah 2013
El Mreiti,Atar Group, West Africa	F4-107-1	1100	13.0	Gilleaudeau and Kah 2013
El Mreiti,Atar Group, West Africa	F4-108-1	1100	6.8	Gilleaudeau and Kah 2013
El Mreiti,Atar Group, West Africa	F4-109-1	1100	6.1	Gilleaudeau and Kah 2013
El Mreiti,Atar Group, West Africa	F4-113-1	1100	13.1	Gilleaudeau and Kah 2013
El Mreiti,Atar Group, West Africa	F4-114-1	1100	7.2	Gilleaudeau and Kah 2013
El Mreiti,Atar Group, West Africa	F4-115-1	1100	7.8	Gilleaudeau and Kah 2013
El Mreiti,Atar Group, West Africa	F4-116-1	1100	7.8	Gilleaudeau and Kah 2013
El Mreiti,Atar Group, West Africa	F4-117-1	1100	7.9	Gilleaudeau and Kah 2013
Atar Group, West Africa	ATS-31	1100	14.8	Kah et al. 2012
Atar Group, West Africa	ATS-53	1100	92.9	Kah et al. 2012
Atar Group, West Africa	ATS-61	1100	16.5	Kah et al. 2012
Atar Group, West Africa	ATS-5-1	1100	13.8	Kah et al. 2012
Atar Group, West Africa	ATS-119	1100	11.6	Kah et al. 2012
Atar Group, West Africa	ATS-154	1100	12.8	Kah et al. 2012
Atar Group, West Africa	ATD-17	1100	18.4	Kah et al. 2012
Atar Group, West Africa	ATD-61	1100	18.2	Kah et al. 2012
Atar Group, West Africa	ATL-51-1	1100	9.3	Kah et al. 2012
Atar Group, West Africa	ATL-58	1100	66.2	Kah et al. 2012
Atar Group, West Africa	ATL-68	1100	4.9	Kah et al. 2012
Atar Group, West Africa	ATL-105	1100	19.3	Kah et al. 2012
Atar Group, West Africa	ATL-110-1	1100	6.3	Kah et al. 2012
Atar Group, West Africa	ATD-27.5	1100	12.3	Kah et al. 2012
Atar Group, West Africa	ATD-45	1100	18.4	Kah et al. 2012
Atar Group, West Africa	R1-Δ-1	1100	16.4	Manning-Berg and Kah in prep
Atar Group, West Africa	R1-Δ-7	1100	25.5	Manning-Berg and Kah in prep
Atar Group, West Africa	R1-Δ-13	1100	14.9	Manning-Berg and Kah in prep
Atar Group, West Africa	R1-Δ-25	1100	12.7	Manning-Berg and Kah in prep
Atar Group, West Africa	R1-Δ-29	1100	12.0	Manning-Berg and Kah in prep
Atar Group, West Africa	R1-Δ-30	1100	10.3	Manning-Berg and Kah in prep
Atar Group, West Africa	R1-Δ-31	1100	5.8	Manning-Berg and Kah in prep
Atar Group, West Africa	R1-Δ-38	1100	12.2	Manning-Berg and Kah in prep
Atar Group, West Africa	R1-Δ-39	1100	6.5	Manning-Berg and Kah in prep
Atar Group, West Africa	R1-Δ-41	1100	9.6	Manning-Berg and Kah in prep
Atar Group, West Africa	R1-Δ-42	1100	6.7	Manning-Berg and Kah in prep
Atar Group, West Africa	R1-Δ-48	1100	8.1	Manning-Berg and Kah in prep
Atar Group, West Africa	R1-Δ-54-1	1100	9.6	Manning-Berg and Kah in prep
Atar Group, West Africa	R1-Δ-55	1100	7.1	Manning-Berg and Kah in prep
Atar Group, West Africa	R1-Δ-58	1100	7.6	Manning-Berg and Kah in prep
Atar Group, West Africa	R1-Δ-60	1100	4.7	Manning-Berg and Kah in prep
Atar Group, West Africa	R1-Δ-61	1100	4.6	Manning-Berg and Kah in prep
Atar Group, West Africa	R1-Δ-65	1100	9.9	Manning-Berg and Kah in prep
Atar Group, West Africa	R1-Δ-68-1	1100	18.3	Manning-Berg and Kah in prep
Atar Group, West Africa	R1-Δ-72	1100	6.6	Manning-Berg and Kah in prep
Atar Group, West Africa	R1-Δ-76	1100	12.4	Manning-Berg and Kah in prep
Atar Group, West Africa	R1-Δ-79-1	1100	8.3	Manning-Berg and Kah in prep
Atar Group, West Africa	R1-Δ-81-1	1100	5.2	Manning-Berg and Kah in prep
Atar Group, West Africa	R1-Δ-85	1100	6.0	Manning-Berg and Kah in prep
Atar Group, West Africa	R1-Δ-89	1100	16.8	Manning-Berg and Kah in prep
Atar Group, West Africa	R1-Δ-90	1100	4.4	Manning-Berg and Kah in prep
Atar Group, West Africa	R1-Δ-97	1100	3.1	Manning-Berg and Kah in prep
Atar Group, West Africa	R1-Δ-98	1100	9.2	Manning-Berg and Kah in prep

Atar Group, West Africa	R1-Δ-100	1100	3.6	Manning-Berg and Kah in prep
Atar Group, West Africa	R1-Δ-103	1100	3.1	Manning-Berg and Kah in prep
Atar Group, West Africa	R1-Δ-106	1100	2.2	Manning-Berg and Kah in prep
Chattisgarh Supergroup, India	SRJ-2I	1000	25.9	Bickford et al. 2011
Chattisgarh Supergroup, India	SRJ-3I	1000	22.8	Bickford et al. 2011
Chattisgarh Supergroup, India	SRJ-4I	1000	13.0	Bickford et al. 2011
Chattisgarh Supergroup, India	SRJ-6I	1000	20.6	Bickford et al. 2011
Chattisgarh Supergroup, India	SRJ-8I	1000	37.8	Bickford et al. 2011
Chattisgarh Supergroup, India	SRJ-9I	1000	85.2	Bickford et al. 2011
Chattisgarh Supergroup, India	TML-3I	1000	7.7	Bickford et al. 2011
Chattisgarh Supergroup, India	TML-6I	1000	17.8	Bickford et al. 2011
Chattisgarh Supergroup, India	TML-7I	1000	14.9	Bickford et al. 2011
Sukhaya Tunguska Formation, Russia	AN59	950	29.5	Sergeev et al. 1997
Sukhaya Tunguska Formation, Russia	AN60	950	46.1	Sergeev et al. 1997
Sukhaya Tunguska Formation, Russia	AN61	950	25.7	Sergeev et al. 1997
Sukhaya Tunguska Formation, Russia	AN62	950	76.4	Sergeev et al. 1997
Sukhaya Tunguska Formation, Russia	AN63	950	110.5	Sergeev et al. 1997
Sukhaya Tunguska Formation, Russia	AN64	950	17.0	Sergeev et al. 1997
Sukhaya Tunguska Formation, Russia	AN66	950	28.1	Sergeev et al. 1997
Sukhaya Tunguska Formation, Russia	AN67	950	20.0	Sergeev et al. 1997
Sukhaya Tunguska Formation, Russia	AN68	950	20.2	Sergeev et al. 1997
Akademikerbeen Group, Spitsbergen	AN1	775	29.2	Knoll and Swett 1990
Akademikerbeen Group, Spitsbergen	AN2	775	40.3	Knoll and Swett 1990
Akademikerbeen Group, Spitsbergen	AN3	775	17.3	Knoll and Swett 1990
Akademikerbeen Group, Spitsbergen	AN4	775	6.2	Knoll and Swett 1990
Akademikerbeen Group, Spitsbergen	AN5	775	63.2	Knoll and Swett 1990
Akademikerbeen Group, Spitsbergen	AN6	775	65.0	Knoll and Swett 1990
Akademikerbeen Group, Spitsbergen	AN7	775	27.6	Knoll and Swett 1990
Akademikerbeen Group, Spitsbergen	AN8	775	74.3	Knoll and Swett 1990
Akademikerbeen Group, Spitsbergen	AN9	775	95.4	Knoll and Swett 1990
Akademikerbeen Group, Spitsbergen	AN10	775	59.4	Knoll and Swett 1990
Akademikerbeen Group, Spitsbergen	AN12	775	71.5	Knoll and Swett 1990
Akademikerbeen Group, Spitsbergen	AN13	775	15.8	Knoll and Swett 1990
Akademikerbeen Group, Spitsbergen	AN14	775	49.4	Knoll and Swett 1990
Akademikerbeen Group, Spitsbergen	AN16	775	15.1	Knoll and Swett 1990
Akademikerbeen Group, Spitsbergen	AN18	775	17.6	Knoll and Swett 1990
Akademikerbeen Group, Spitsbergen	AN19	775	9.4	Knoll and Swett 1990
Akademikerbeen Group, Spitsbergen	AN20	775	34.4	Knoll and Swett 1990
Akademikerbeen Group, Spitsbergen	AN21	775	276.2	Knoll and Swett 1990
Akademikerbeen Group, Spitsbergen	AN22	775	34.6	Knoll and Swett 1990
Limestone-Dolomite Series, East Greenland	AN24	775	126.1	Knoll et al. 1986
Limestone-Dolomite Series, East Greenland	AN25	775	59.3	Knoll et al. 1986
Limestone-Dolomite Series, East Greenland	AN26	775	45.7	Knoll et al. 1986
Limestone-Dolomite Series, East Greenland	AN28	775	5.3	Knoll et al. 1986
Limestone-Dolomite Series, East Greenland	AN29	775	61.2	Knoll et al. 1986
Limestone-Dolomite Series, East Greenland	AN30	775	307.6	Knoll et al. 1986
Limestone-Dolomite Series, East Greenland	AN31	775	310.7	Knoll et al. 1986
Limestone-Dolomite Series, East Greenland	AN32	775	38.3	Knoll et al. 1986
Limestone-Dolomite Series, East Greenland	AN33	775	33.3	Knoll et al. 1986
Limestone-Dolomite Series, East Greenland	AN34	775	15.5	Knoll et al. 1986
Limestone-Dolomite Series, East Greenland	AN35	775	8.3	Knoll et al. 1986
Limestone-Dolomite Series, East Greenland	AN36	775	102.9	Knoll et al. 1986
Limestone-Dolomite Series, East Greenland	AN37	775	67.0	Knoll et al. 1986
Limestone-Dolomite Series, East Greenland	AN38	775	41.4	Knoll et al. 1986

Shaler Group, Arctic Canada	AN41	775	5.6	Jones et al. 2010
Shaler Group, Arctic Canada	AN42	775	17.3	Jones et al. 2010
Shaler Group, Arctic Canada	AN43	775	45.7	Jones et al. 2010
Shaler Group, Arctic Canada	AN44	775	7.0	Jones et al. 2010
Shaler Group, Arctic Canada	AN45	775	154.8	Jones et al. 2010
Shaler Group, Arctic Canada	AN46	775	6.2	Jones et al. 2010
Shaler Group, Arctic Canada	AN47	775	40.3	Jones et al. 2010
Shaler Group, Arctic Canada	AN48	775	14.8	Jones et al. 2010
Shaler Group, Arctic Canada	AN49	775	116.9	Jones et al. 2010
Shaler Group, Arctic Canada	AN50	775	8.4	Jones et al. 2010
Shaler Group, Arctic Canada	AN51	775	10.9	Jones et al. 2010
Shaler Group, Arctic Canada	AN52	775	57.5	Jones et al. 2010
Shaler Group, Arctic Canada	AN53	775	6.9	Jones et al. 2010
Shaler Group, Arctic Canada	AN54	775	6.0	Jones et al. 2010
Shaler Group, Arctic Canada	AN55	775	12.8	Jones et al. 2010
Shaler Group, Arctic Canada	AN56	775	17.8	Jones et al. 2010
Lagoa Do Jacare Formation, Brazil	KM7-14.0.0	650	38.1	Misi, A. et al., 2007
Lagoa Do Jacare Formation, Brazil	KM7-14-01.0	650	21.9	Misi, A. et al., 2007
Lagoa Do Jacare Formation, Brazil	KM7-14-02.0	650	49.0	Misi, A. et al., 2007
Lagoa Do Jacare Formation, Brazil	KM7-14-03.0	650	17.6	Misi, A. et al., 2007
Lagoa Do Jacare Formation, Brazil	KM7-14-04.0	650	12.5	Misi, A. et al., 2007
Lagoa Do Jacare Formation, Brazil	KM7-14-05.0	650	22.0	Misi, A. et al., 2007
Lagoa Do Jacare Formation, Brazil	KM7-14-06.0	650	38.9	Misi, A. et al., 2007
Lagoa Do Jacare Formation, Brazil	KM7-14-07.0	650	54.1	Misi, A. et al., 2007
Lagoa Do Jacare Formation, Brazil	KM7-14-08.0	650	32.8	Misi, A. et al., 2007
Lagoa Do Jacare Formation, Brazil	KM7-14-09.0	650	23.8	Misi, A. et al., 2007
Huttenburg Fm, Namibia	S86A-971.2	650	105.2	Kaufman et al., 2009
Huttenburg Fm, Namibia	S86A-976.0	650	77.5	Kaufman et al., 2009
Huttenburg Fm, Namibia	S86A-977.0	650	120.9	Kaufman et al., 2009
Huttenburg Fm, Namibia	S86A-980.8	650	53.5	Kaufman et al., 2009
Huttenburg Fm, Namibia	S86A-985.1	650	147.1	Kaufman et al., 2009
Huttenburg Fm, Namibia	S86A-987.8	650	28.4	Kaufman et al., 2009
Huttenburg Fm, Namibia	S86A-988.2	650	41.8	Kaufman et al., 2009
Huttenburg Fm, Namibia	S86A-1033.8	650	424.1	Kaufman et al., 2009
Huttenburg Fm, Namibia	S86A-1060.2	650	156.2	Kaufman et al., 2009
Huttenburg Fm, Namibia	S86A-1077.1	650	57.4	Kaufman et al., 2009
Huttenburg Fm, Namibia	S86A-1144.8	650	52.8	Kaufman et al., 2009
Huttenburg Fm, Namibia	S86A-1145.1	650	47.7	Kaufman et al., 2009
Huttenburg Fm, Namibia	S86A-1148.4	650	55.6	Kaufman et al., 2009
Huttenburg Fm, Namibia	S86A-1213.2	650	51.1	Kaufman et al., 2009
Dhaiqa Fm, NW Arabian shield, Saudi Arabia	Dhaiqa-7	600	61.0	Miller et al. 2008
Dhaiqa Fm, NW Arabian shield, Saudi Arabia	Dhaiqa-21	600	62.5	Miller et al. 2008
Dhaiqa Fm, NW Arabian shield, Saudi Arabia	Dhaiqa-26	600	30.7	Miller et al. 2008
Dhaiqa Fm, NW Arabian shield, Saudi Arabia	Dhaiqa-34	600	37.5	Miller et al. 2008
Dhaiqa Fm, NW Arabian shield, Saudi Arabia	Dhaiqa-38	600	19.8	Miller et al. 2008
Dhaiqa Fm, NW Arabian shield, Saudi Arabia	Dhaiqa-38b	600	40.1	Miller et al. 2008
Dhaiqa Fm, NW Arabian shield, Saudi Arabia	Dhaiqa-39	600	38.7	Miller et al. 2008
Dhaiqa Fm, NW Arabian shield, Saudi Arabia	Dhaiqa-46	600	24.9	Miller et al. 2008
Dhaiqa Fm, NW Arabian shield, Saudi Arabia	Dhaiqa-49	600	59.0	Miller et al. 2008
Dhaiqa Fm, NW Arabian shield, Saudi Arabia	Dhaiqa-51	600	19.8	Miller et al. 2008
Dhaiqa Fm, NW Arabian shield, Saudi Arabia	Dhaiqa-54	600	24.9	Miller et al. 2008
Dhaiqa Fm, NW Arabian shield, Saudi Arabia	M1-with fossil	600	62.2	Miller et al. 2008
Dhaiqa Fm, NW Arabian shield, Saudi Arabia	N-2-3	600	113.4	Miller et al. 2008
Dhaiqa Fm, NW Arabian shield, Saudi Arabia	N-2-11	600	73.1	Miller et al. 2008

Dhaiqa Fm, NW Arabian shield, Saudi Arabia	N-2-16	600	68.9	Miller et al. 2008
Yangjiaping, Doushantuo Fm, South China	YD-01	551	12.9	Cui et al. 2015
Yangjiaping, Doushantuo Fm, South China	YD-02	551	47.1	Cui et al. 2015
Yangjiaping, Doushantuo Fm, South China	YD-03	551	11.7	Cui et al. 2015
Yangjiaping, Doushantuo Fm, South China	YD-04	551	25.1	Cui et al. 2015
Yangjiaping, Doushantuo Fm, South China	YD-05	551	71.0	Cui et al. 2015
Yangjiaping, Doushantuo Fm, South China	YD-06	551	13.9	Cui et al. 2015
Yangjiaping, Doushantuo Fm, South China	YD-07	551	28.4	Cui et al. 2015
Yangjiaping, Doushantuo Fm, South China	YD-08	551	6.6	Cui et al. 2015
Yangjiaping, Doushantuo Fm, South China	YD-09	551	6.2	Cui et al. 2015
Yangjiaping, Doushantuo Fm, South China	YD-10	551	26.9	Cui et al. 2015
Yangjiaping, Doushantuo Fm, South China	YD-11	551	50.4	Cui et al. 2015
Yangjiaping, Doushantuo Fm, South China	YD-12	551	25.1	Cui et al. 2015
Yangjiaping, Doushantuo Fm, South China	YD-13	551	23.7	Cui et al. 2015
Yangjiaping, Doushantuo Fm, South China	YD-14	551	55.0	Cui et al. 2015
Yangjiaping, Doushantuo Fm, South China	YD-15	551	115.4	Cui et al. 2015
Yangjiaping, Doushantuo Fm, South China	YD-16	551	221.5	Cui et al. 2015
Yangjiaping, Doushantuo Fm, South China	YD-17	551	253.8	Cui et al. 2015
Yangjiaping, Doushantuo Fm, South China	YD-18	551	195.9	Cui et al. 2015
Yangjiaping, Doushantuo Fm, South China	YD-19	551	101.2	Cui et al. 2015
Yangjiaping, Doushantuo Fm, South China	YD-20	551	14.5	Cui et al. 2015
Yangjiaping, Doushantuo Fm, South China	YD-21	551	10.9	Cui et al. 2015
Yangjiaping, Doushantuo Fm, South China	YD-22	551	29.1	Cui et al. 2015
Yangjiaping, Doushantuo Fm, South China	YD-23	551	40.1	Cui et al. 2015
Orthoceras limestone, Ohio	BH21	465	27.5	Hazen collection
Bryozoa limestone, Ohio	BH24	465	7.2	Hazen collection
Branch Hill limestone, Ohio	BH25	465	10.9	Hazen collection
Jersey Shore Station, Pennsylvania	BH26	465	7.0	Hazen collection
Madison County, Kentucky	BH31	465	14.3	Hazen collection
Ripley, Ohio	BH36	465	8.3	Hazen collection
Washington, Kentucky	BH37	465	11.4	Hazen collection
Bryozoa limestone, Ohio	BH33	465	3.5	Hazen collection
Coburn Fm, Pennsylvania	BH41	465	10.9	Hazen collection
La Silla Fm, Argentina	AF-23	464	22.9	Thompson and Kah 2012
La Silla Fm, Argentina	AF-30	464	12.4	Thompson and Kah 2012
La Silla Fm, Argentina	SJF08-14	464	16.1	Thompson and Kah 2012
La Silla Fm, Argentina	LFG-27	464	40.5	Thompson and Kah 2012
La Silla Fm, Argentina	LFG-55	464	17.6	Thompson and Kah 2012
La Silla Fm, Argentina	LS-01	464	29.7	Thompson and Kah 2012
La Silla Fm, Argentina	SJC-116	464	42.1	Thompson and Kah 2012
West Newfoundland	TH-1	464	15.3	Thompson and Kah 2012
West Newfoundland	TH-18	464	41.7	Thompson and Kah 2012
Ludlow Fm, Silurian	BH28	432	30.3	Hazen collection
Clay's Ferry Fm, Kentucky	BH49	457	7.1	Hazen collection
Limestone Quarry, Dickensonville, Virginia	BH29	406	7.7	Hazen collection
John Boyd Thatchen State Park, New York	BH30	406	12.8	Hazen collection
Cincinnati, Ohio	BH32	451	6.5	Hazen collection
Richmond, Indiana	BH27	451	7.3	Hazen collection
Blue Stone limestone Quarry, New York	BH34	406	19.0	Hazen collection
Isle La Motte, Vermont	BH38	406	21.6	Hazen collection
Isle La Motte, Vermont	BH39	406	12.2	Hazen collection
Isle La Motte, Vermont	BH40	406	22.9	Hazen collection
Tenby, Wales	BH43	329	69.0	Hazen collection
Plainville Quarry, Ohio	BH44	388	17.4	Hazen collection

Everett Quarry, Missouri	BH45	311	15.4	Hazen collection
Cheddar Gorge, England	BH46	329	65.5	Hazen collection
Mammoth Cave National Park, Kentucky	BH47	341	190.6	Hazen collection
Sheep Mountain	BH51	341	53.6	Hazen collection
Limestone Quarry, Ohio	BH52	388	4.2	Hazen collection
Hamilton Group, New York	BH22	388	3.6	Hazen collection
Madison limestone, Garrett Co., Maryland	BH23	341	31.7	Hazen collection
Coral limestone, Michigan	BH53	389	21.0	Hazen collection
Lydstep Haven coral, Wales	BH35	329	70.4	Hazen collection
Shark Bay, Australia	BH20	0	28.4	Hazen collection

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## ***SI 2. Data filtering***

Diagenetic alteration can be problematic when interpreting the chemical composition carbonate rocks; therefore, we carefully screened carbonate specimens for a range of diagenetic effects and hydrothermal alteration by combining geologic, petrographic, and element and isotope geochemical analyses. For the published literature, we compiled data only for samples that are considered to reflect primary depositional environments. For our own analyses, we selected samples from within known sedimentological and stratigraphic context, most of which are from previously investigated rock units. Petrographic analysis was used to selected sample areas preserving original sedimentary fabrics, which suggest the least interaction with diagenetic fluids, and samples were microdrilled from these localities. To identify possible affects of diagenesis, we screened the samples based on the classic geochemical tracers of late diagenesis, including major and minor elements (e.g., Fe, Mn, Sr) as well as C and O isotopic signatures. We did not fix the selection criteria; rather, we adopted the selection criteria used by individual researchers for the different localities and published in their original descriptions. Readers can refer to the original papers that describe the samples in the reference lists below ([SI 6. References](#)). In addition, we studied REE patterns and Eu anomalies to avoid samples with significant hydrothermal alternation (e.g., Frimmel, 2009). Both Zn and Fe partition coefficient ( $K_d$ ) from fluid to carbonates increase with increasing diagenesis as shown by earlier work of Brand and Veizer (1980). Also,  $K_d$  (Fe) increases faster compared to  $K_d$  (Zn), from 1 to 20 and from 5.2 to 5.5 for Fe and Zn, respectively. So, diagenesis will cause decrease in Zn/Fe ratios by incorporating more Fe than Zn in carbonates. Even though, every carbonates we studied today have been through diagenesis. That is why we are adopting a statistic treatment of our global data.

### ***SI 3. Statistical analysis of data***

Temporal data sets are subject to biases associated with sampling: recent geological eras are commonly represented by more samples, and some formations are represented by multiple analyses. Since we do not know exactly how the effect of diagenesis, local primary production, mineralogy and kinetic effect influence on the Zn/Fe values, we therefore adopted three discrete approaches to statistically evaluate Zn/Fe data through time aiming to investigate the population behavior. First, we divided the entire sample population into eight bins of different duration to make sure each bin has statistically meaningful sample numbers (where  $n > 50$ , expect for one, bin with  $n = 38$ ). The bins were chosen to reflect current hypotheses for  $pO_2$  evolution through time; specifically, we chose a bin boundary at 800 Ma to reflect the recent hypothesis of Planavsky et al. (2014) concerning mid-Neoproterozoic oxygenation, and we broke out Cryogenian, Ediacaran and Lower Paleozoic bins in an attempt to illuminate existing hypothesis concerning Ediacaran-Cambrian oxygen increase. Each bin contains samples from at least two different geological formations. The choice of the bin breaking points is based on convenience and what is already know about  $pO_2$  evolution through Earth history. For example, the first bin is from 3.5 – 2.5 Ga, which is the Pre-GOE time, where we expect low  $pO_2$  and we observe low Zn/Fe and little variation of Zn/Fe. We divide the following bin into 2.5-2.0 Ga and 2.0-1.5 Ga, to have statistical meaning number of samples in each bin. Then we have three different formation from ~800 Ma with high Zn/Fe compared to Pre-1.0Ga samples and therefore we make a bin from 1.5-0.8 Ga. For the following time periods, we divide bins into several geological meaningful groups, as 800-635 Ma (Cryogenian), 635-541 Ma (Ediacaran), 541-300 Ma (earlier

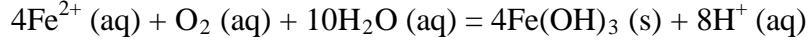


Paleozoic), and 300-0 Ma (later Paleozoic, Mesozoic and Cenozoic) to make self-consistent and statistically meaningful sample subsets.

We then performed a box-whisker plot for all data (Fig. S3), where median, 50%, and outliers (outside of 3 sigma of the population) of Zn/Fe values were calculated for each bin, which are shown with red lines (orange lines in Fig. 2), blue boxes, and red crosses, respectively in Fig. S3. In a second approach, we divided data into the same ten bins as in the previous approach, but we plotted histograms for each bin (Fig. S4). Data in nearly all of the bins follows a lognormal distribution, which permits calculation of means and standard deviations for each of the 10 bins (shown with blue lines in Fig. 2). In a third approach, we calculated average composition for each geological formation, which reduces the influence from sampling bias. For example, some localities/formations are represented by 20 samples, whereas others incorporate 10 or fewer. The formation averages are shown in Fig. 2 and a polynomial fit is calculated and displayed in Fig. S5.

#### SI 4. Assumptions, deviation of equations and sensitivity tests

Fe oxidization chemical reaction can be written as:



$$K = \frac{a_{(\text{H}^+)}^8}{a_{(\text{H}_2\text{O})}^{10} \times a_{(\text{Fe}^{2+})}^4 \times a_{(\text{O}_2)}} \quad [1],$$

where  $K$  is equilibrium constant and  $\alpha$  is activity. In this equation, we assume that when  $\text{Fe}^{2+}$  oxidizes to  $\text{Fe}^{3+}$  and is precipitated from the aqueous system as iron hydroxide  $[\text{Fe}(\text{OH})_3]$ , thus only  $\text{Fe}^{2+}$  is incorporated into carbonates. Assuming equilibrium between atmosphere and surface ocean on hundredmillion year time scales (the smallest bin size is around 100 million years, Bin6: 635-541 Ma), we replace  $\alpha(\text{O}_2)$  in equation [1] with oxygen fugacity in the atmosphere,  $f_{(\text{O}_2)}$ , as

$$K = \frac{a_{(\text{H}^+)}^8}{a_{(\text{H}_2\text{O})}^{10} \times a_{(\text{Fe}^{2+})}^4 \times f_{(\text{O}_2)}} \quad [2].$$

Then we can reorganize equation [2] to

$$\log K = 4\log \frac{1}{a_{(\text{Fe}^{2+})}} - \log f_{(\text{O}_2)} + \log \frac{a_{(\text{H}^+)}^8}{a_{(\text{H}_2\text{O})}^{10}} \quad [3],$$

if we assume the Zn concentrations in seawater and partitioning of Zn/Fe from seawater to carbonate minerals are constant over Earth's history. Therefore, we can write the equation normalized to  $\text{Zn}^{2+}$  as

$$\log K = 4\log a_{(\text{Zn}^{2+}/\text{Fe}^{2+})}^P - \log f_{(\text{O}_2)}^P + \log \frac{a_{(\text{H}^+)}^8}{a_{(\text{H}_2\text{O})}^{10}}$$

$$\log K = 4 \log a_{(Zn^{2+}/Fe^{2+})}^M - \log f_{(O_2)}^M + \log \frac{a_{(H^+)}^8}{a_{(H_2O)}^{10}} \quad [4],$$

in which superscripts P and M indicate the past and modern parameters. Assuming pH and K are constant, we can simplify the relationship between Fe/Zn ratios and  $f_{(O_2)}$ , as  $f_{(O_2)}^P = r^4 \cdot f_{(O_2)}^M$  [6], where  $f_{(O_2)}^P$  is the oxygen fugacity in the past (any time in Earth's history),  $f_{(O_2)}^M$  is the oxygen fugacity in modern time, and

$$r = \frac{(Fe^{2+}/Zn^{2+})^M}{(Fe^{2+}/Zn^{2+})^P} = \frac{(Zn^{2+}/Fe^{2+})^P}{(Zn^{2+}/Fe^{2+})^M} \quad [6],$$

provides Zn/Fe ratios in past carbonate normalized to the modern values. If we assume that atmospheric  $O_2$  is in equilibrium with the shallow marine environment, and that we know the current atmospheric  $P_{(O_2)}$  (0.21) and the modern seawater Zn/Fe ratios as reflected in Zn/Fe ratios of marine carbonates, we can therefore calculate  $f_{(O_2)}$  at any given time of Earth's history if we know the Zn/Fe ratio.

### ***Discussion of assumptions***

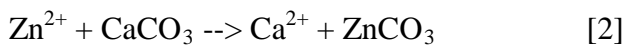
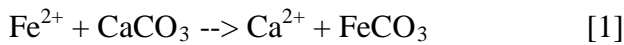
This approach requires several critical assumptions, one of which is that we assume the pH of the ocean has not changed significantly through Earth's history—an assumption supported by some studies (Grotzinger and Kasting, 1993; Sumner and Grotzinger, 2004). However, we do recognize the possible great influence of pH change on quantitative constraint of atmospheric  $O_2$ . We also assume a constant Zn concentration in seawater through time. Although Zn concentrations, theoretically may have responded to changes in enzymatic use during biosphere evolution, there is no evidence for change in Zn bioavailability change through time (Robbins et al., 2013; Scott et al., 2013). However, Zn

concentrations in seawater may drop during the Mid-Proterozoic due to some extent of euxinia (~1~10 % of modern seafloor area) (Reinhard et al., 2013). These two important factors (pH and Zn content) may contribute on various degrees to the accuracy of atmospheric O<sub>2</sub> estimation in our model. However, we do not have a quantitative understanding of either at present. Therefore, we develop a first order quantification of secular evolution of atmospheric O<sub>2</sub> with the assumptions that neither pH nor Zn content in seawater change significantly during the time intervals we investigated.

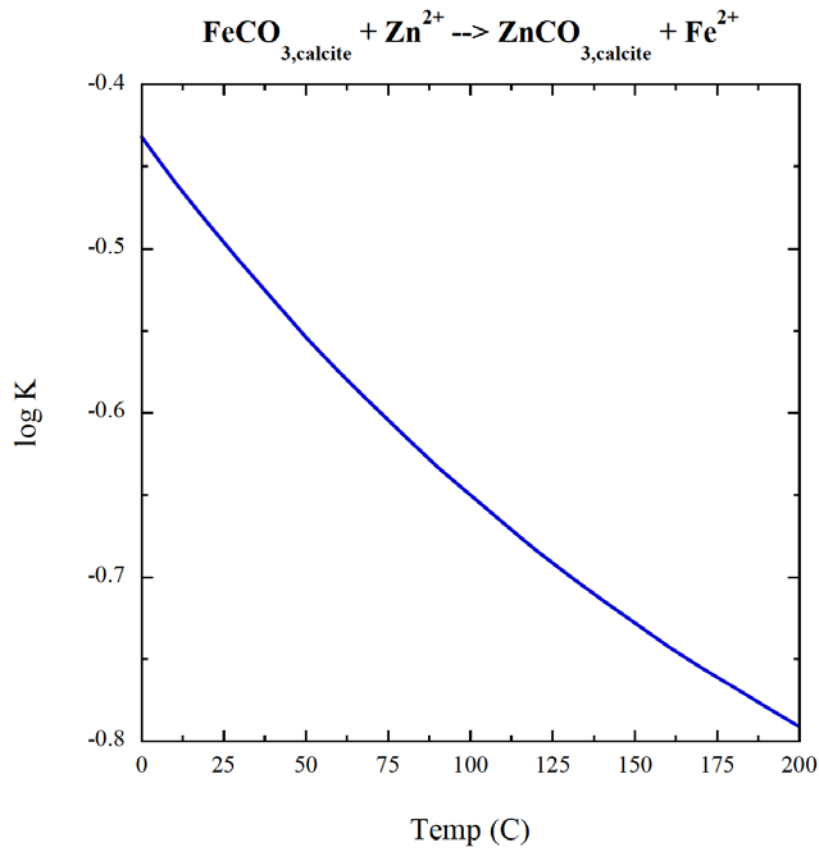
### ***Sensitivity tests of the modeling***

Here, we evaluate how temperature and pH changes influence the modeling results in seawater through Earth's history.

First, we perform a sensitivity test that assuming three different ***temperatures***:



By combining eqn (1) and (2), we have  $\text{FeCO}_3 + \text{Zn}^{2+} \rightarrow \text{ZnCO}_3 + \text{Fe}^{2+}$ , where equilibrium constant K can be written as  $\log K = \log (a(\text{ZnCO}_3)/a(\text{FeCO}_3)) - \log (a(\text{Zn}^{2+})/a(\text{Fe}^{2+}))$ , where a are activities for different components. We then calculate K values at different temperatures using SUPCRT92 (Johnson et al., 1992) and the results are shown in the following figure.



Generally, when T increases, log K (also K) decreases, but still not so much.

For example,

When T = 25 °C, log K = -0.496, K = 0.32;

When T = 60 °C, log K = -0.575, K = 0.27;

When T = 100 °C, log K = -0.65, K = 0.22.

If we take the most extreme example: Assuming modern T is 25°C and Archean T is 100°C.

The  $f\text{O}_2$  estimate in the Archean will increase by approximately 40%. Therefore, even if we assume the most extreme ocean temperature (100°C), the uncertainty caused by temperature

change on  $fO_2$  estimate is much smaller compared to uncertainties generated by Zn/Fe variations in carbonates (Figure 3), which is at least one order of magnitude.

Second, we carry out a sensitivity test on  $pH$  change:

$$\log K = 4 \log a_{\left(\frac{Zn^{2+}}{Fe}\right)}^P - \log f_{(O_2)}^P + \log \frac{a_{(H^+)}^8}{a_{(H_2O)}^{10}}^P$$

$$\log K = 4 \log a_{(Zn^{2+}/Fe^{2+})}^M - \log f_{(O_2)}^M + \log \frac{a_{(H^+)}^8}{a_{(H_2O)}^{10}}^M$$

The modern day seawater has a pH of  $\sim 8$  (M in the equation). If we assume a lower pH in the past (P in the equation), we can rearrange the equations as

$$4 \log a_{(Zn^{2+}/Fe^{2+})}^P - \log f_{(O_2)}^P + \log \frac{a_{(H^+)}^8}{a_{(H_2O)}^{10}}^P = 4 \log a_{(Zn^{2+}/Fe^{2+})}^M - \log f_{(O_2)}^M + \log \frac{a_{(H^+)}^8}{a_{(H_2O)}^{10}}^M$$

$$\frac{a_{(Zn^{2+}/Fe^{2+})}^P{}^4 a_{(H^+)}^P{}^{10}}{a_{(H_2O)}^P{}^{10} f_{(O_2)}^P} = \frac{a_{(Zn^{2+}/Fe^{2+})}^M{}^4 a_{(H^+)}^M{}^{10}}{a_{(H_2O)}^M{}^{10} f_{(O_2)}^M}$$

Final, we get the new pH dependent relationship as

$$f_{(O_2)}^P = \frac{a_{(Zn^{2+}/Fe^{2+})}^P{}^4}{a_{(Zn^{2+}/Fe^{2+})}^M{}^4} \left( \frac{a_{(H^+)}^P}{a_{(H^+)}^M} \right)^8 f_{(O_2)}^M$$

Therefore, if the pH is lower in the past, this will cause the  $O_2$  estimate to be lower. This also means that the  $O_2$  level we estimate is the maximum values assuming the pH was lower in the past. Also, if we change the pH from 8.1 to 7.6, estimated  $O_2$  is four orders of magnitude lower because the  $O_2$  estimate is sensitive pH change. Since we do not know how seawater

pH change through time, and there is no evidence that it changes dramatically on the order of hundreds of million years, we keep the pH constant in our model. However, if we understand better how pH change through time, we can plug in pH in above equations and get a more accurate estimate of  $O_2$  evolution through time.

*SI 5. Figures*

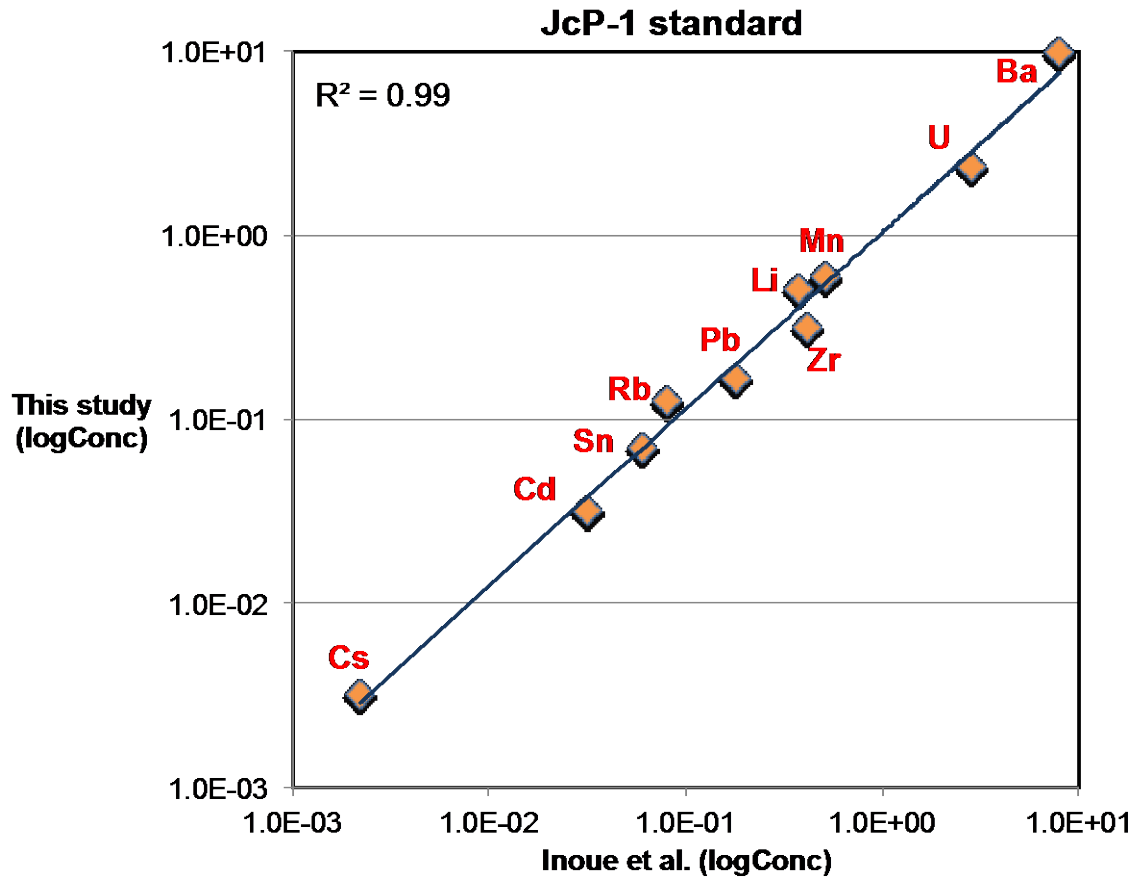


Figure S1. Comparison of measured trace elements with those reported for Inoue et al. (2004) for the JcP-1 standard. The full analytical method is discussed in the method section following the main text.



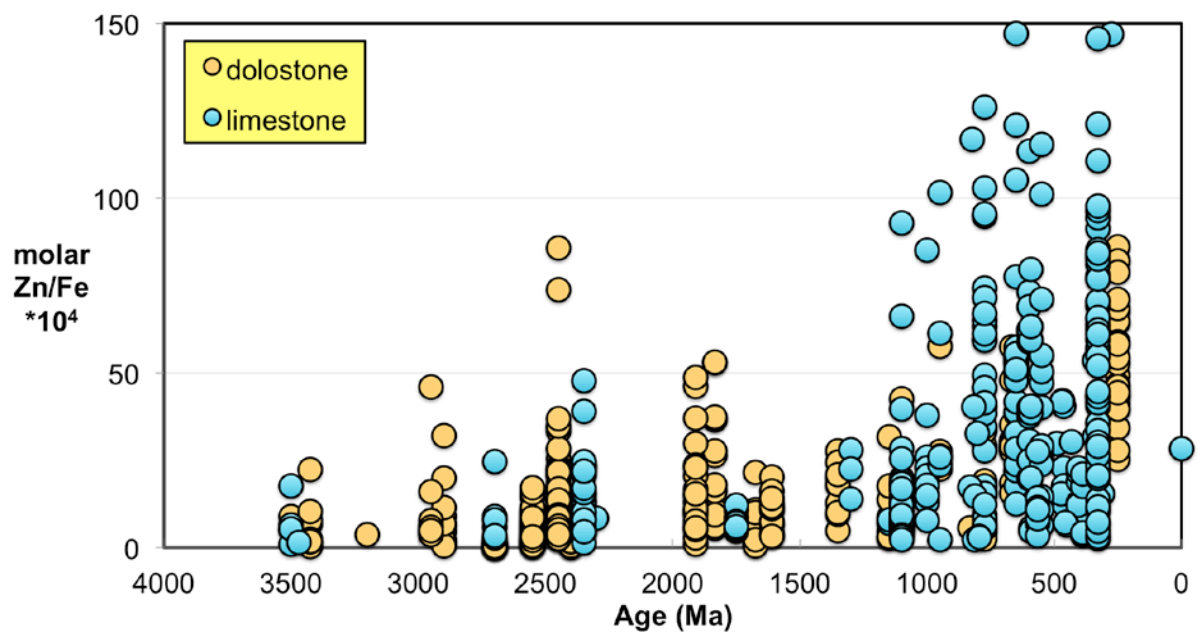


Figure S2. Zn/Fe ratios in marine carbonate, plotted with information on sample lithology.

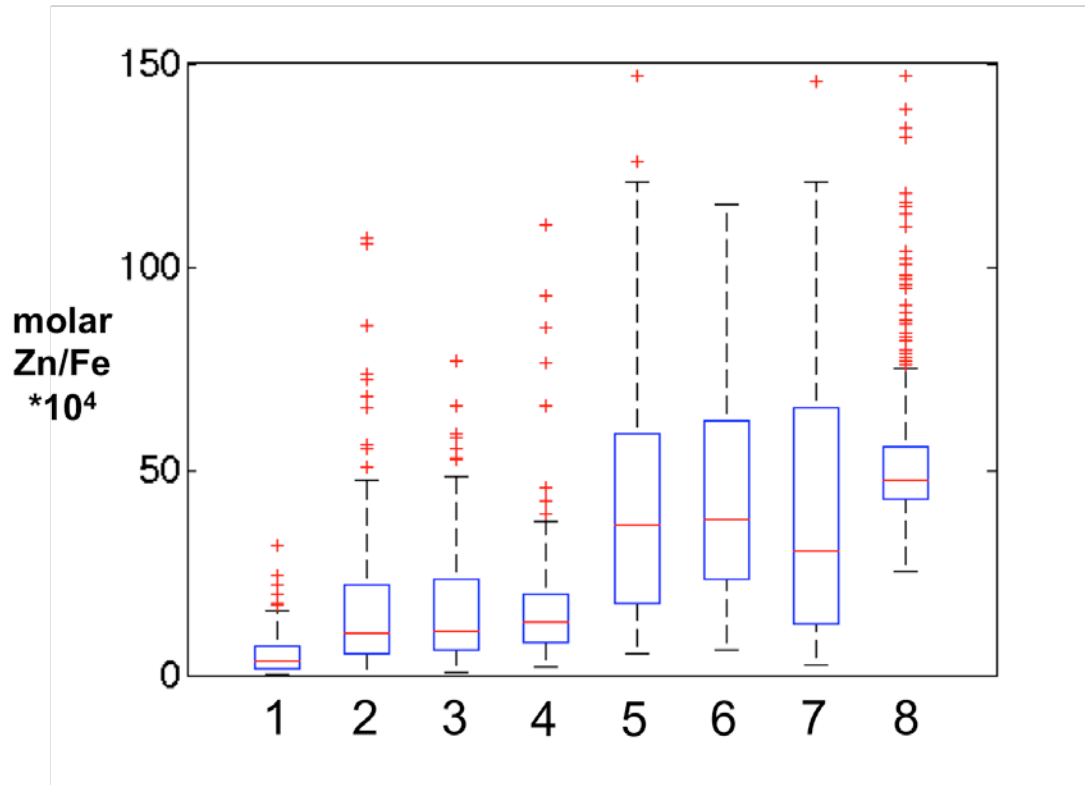
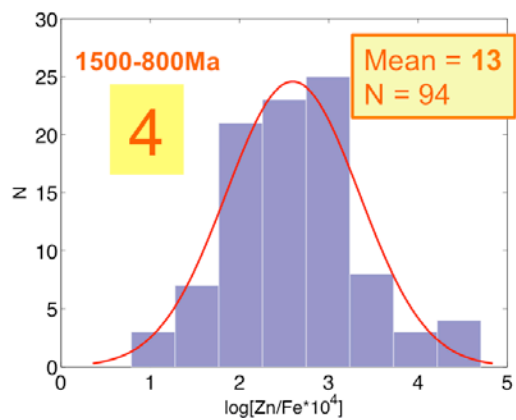
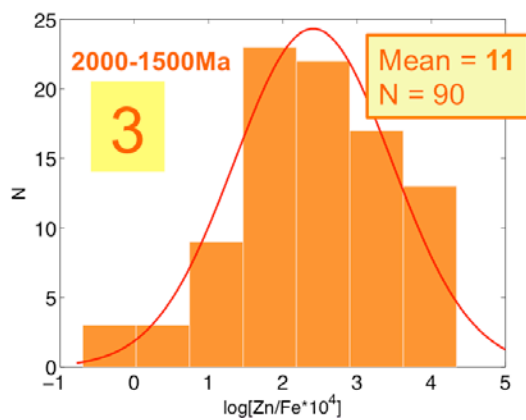
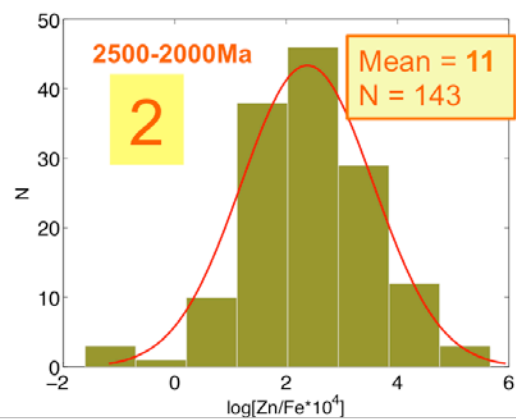
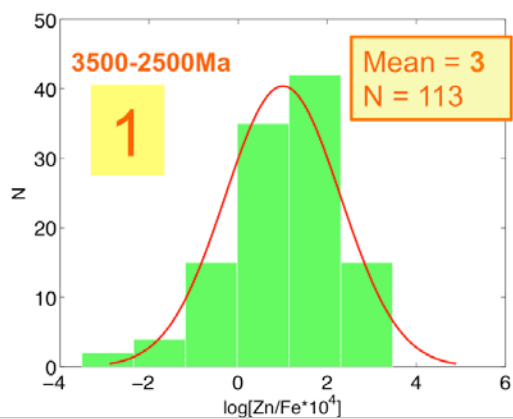


Figure S3. Box-whisker distribution of all samples. The sample population is divided into eight bins (Bin1: 3.5 -2.5 Ga, Bin2: 2.5-2.0 Ga, Bin3: 2.0-1.5 Ga, Bin4: 1.5-0.8Ga, Bin5: 800-635 Ma, Bin6: 635-541Ma, Bin7: 541-300Ma, Bin8: 300-0 Ma) of different duration to make sure each bin has statistically meaningful sample numbers (where  $n > 50$ , expect for one, bin with  $n = 38$ ). Each bin contains samples from at least two different geological formations. We show a Box-whisker plot for each group. Median values are indicated by the red lines and each individual boxes include 50% samples and whiskers mark the 3 sigma boundaries of the group population. Red crosses fall out of whiskers and are considered outliers.



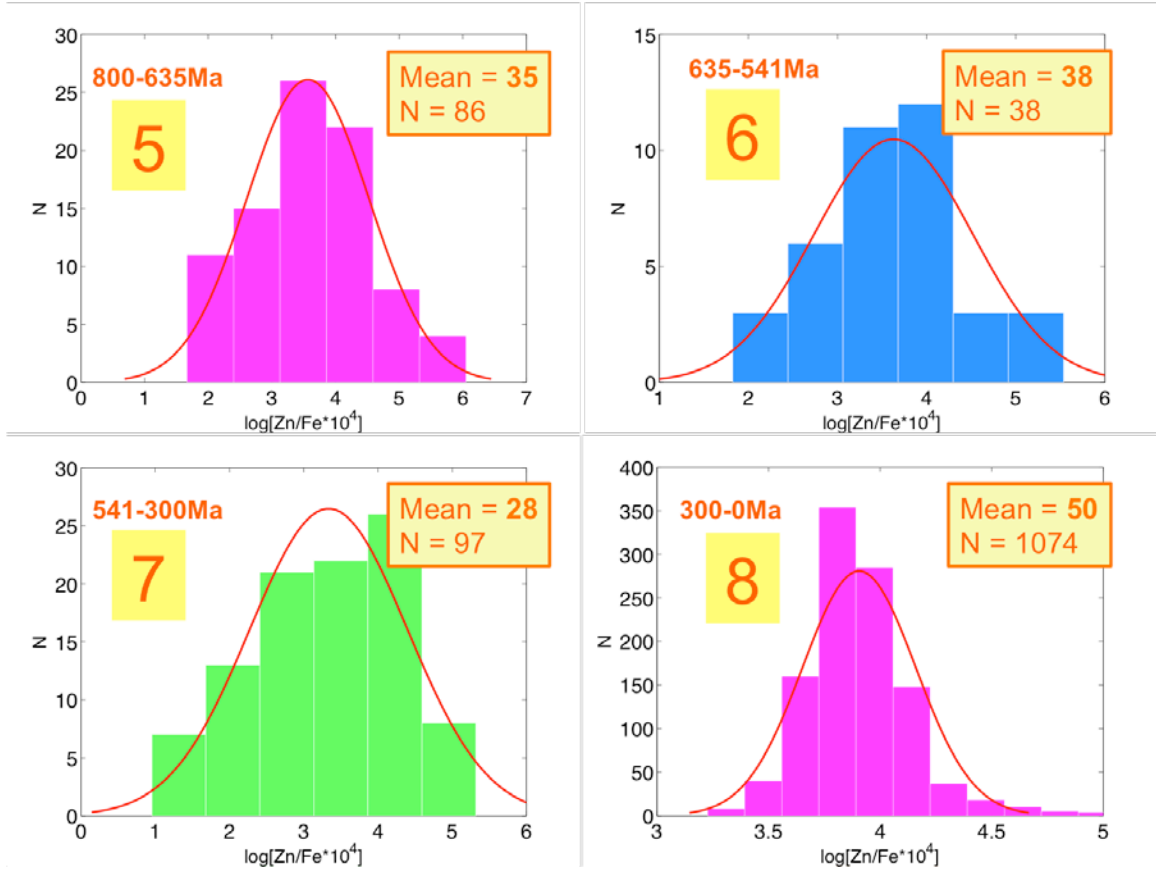


Figure S4. Histograms of Zn/Fe ratios with lognormal fitting in red. We group all data into eight different bins (age distribution of the bins is provided in Fig. S2) and plot the lognormal distribution for each group.

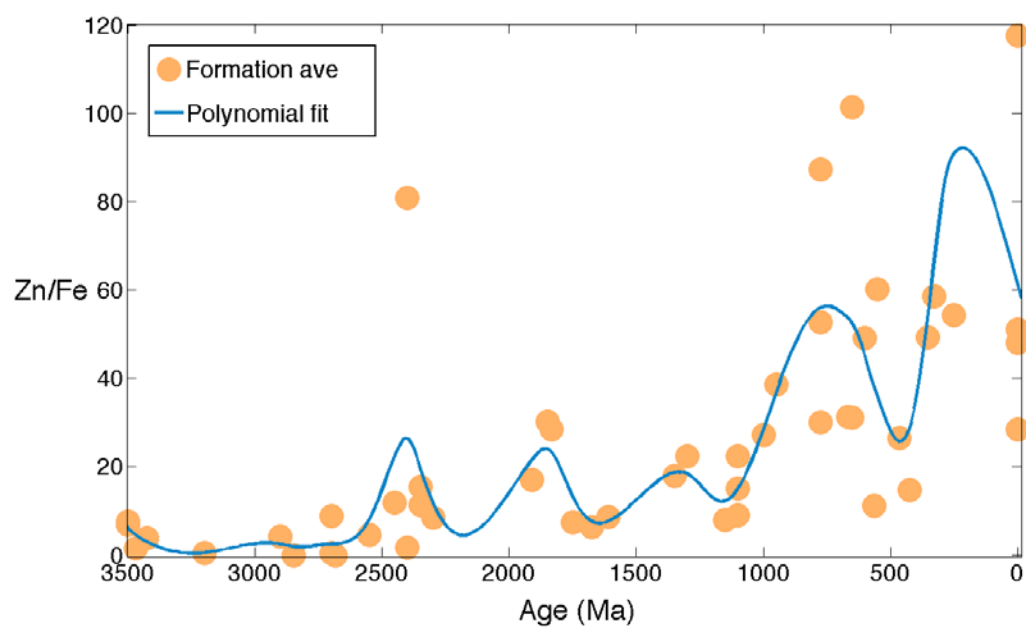


Figure S5. Zn/Fe molar ratio versus time for carbonates averaged by formation. A polynomial fit through the formation average data.

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